# Pendulum (mechanics)

A **pendulum** is a body suspended from a fixed support so that it swings freely back and forth under the influence of gravity. When a pendulum is displaced sideways from its resting, equilibrium position, it is subject to a restoring force due to gravity that will accelerate it back toward the equilibrium position. When released, the restoring force acting on the pendulum's mass causes it to oscillate about the equilibrium position, swinging it back and forth. The mathematics of pendulums are in general quite complicated. Simplifying assumptions can be made, which in the case of a simple pendulum allow the equations of motion to be solved analytically for small-angle oscillations.



Animation of a pendulum showing the velocity and acceleration vectors.

# Simple gravity pendulum

A *simple gravity pendulum*<sup>[1]</sup> is an idealized mathematical model of a real pendulum.<sup>[2][3][4]</sup> This is a weight (or bob) on the end of a massless cord suspended from a pivot, without friction. Since in this model there is no frictional energy loss, when given an initial displacement it will swing back and forth at a constant amplitude. The model is based on these assumptions:

- The rod or cord on which the bob swings is massless, inextensible and always remains taut.
- The bob is a point mass.
- Motion occurs only in two dimensions, i.e. the bob does not trace an ellipse but an arc.
- The motion does not lose energy to friction or air resistance.
- The gravitational field is uniform.

• The support does not move.

The differential equation which represents the motion of a simple pendulum is

(Eq. 1)

where g is the magnitude of the gravitational field,  $\ell$  is the length of the rod or cord, and  $\theta$  is the angle from the vertical to the pendulum.



resolved into components parallel and perpendicular to the bob's instantaneous

motion. The direction of the bob's instantaneous velocity always points along the red axis, which is considered the tangential axis because its direction is always tangent to the circle. Consider Newton's second law,

### F = ma

where F is the sum of forces on the object, m is mass, and a is the acceleration. Because we are only concerned with changes in speed, and because the bob is forced to stay in a circular path, we apply Newton's equation to the tangential axis only. The short violet arrow represents the component of the gravitational force in the tangential axis, and trigonometry can be used to determine its magnitude. Thus,

$$F = -mg\sin heta = ma,$$
 so  $a = -g\sin heta,$ 

where g is the acceleration due to gravity near the surface of the earth. The negative sign on the right hand side implies that  $\theta$  and a always point in opposite directions. This makes sense because when a pendulum swings further to the left, we would expect it to accelerate back toward the right.

This linear acceleration *a* along the red axis can be related to the change in angle  $\theta$  by the arc length formulas; *s* is arc length:

$$egin{aligned} s &= \ell heta, \ v &= rac{ds}{dt} = \ellrac{d heta}{dt}, \ a &= rac{d^2s}{dt^2} = \ellrac{d^2 heta}{dt^2}, \end{aligned}$$

thus:

$$\ell rac{d^2 heta}{dt^2} = -g \sin heta, 
onumber \ rac{d^2 heta}{dt^2} + rac{g}{\ell} \sin heta = 0.$$

### "Torque" derivation of (Eq. 1)

Equation (1) can be obtained using two definitions for torque.

$$oldsymbol{ au} = \mathbf{r} imes \mathbf{F} = rac{d \mathbf{L}}{dt}.$$

First start by defining the torque on the pendulum bob using the force due to gravity.

### $\boldsymbol{ au} = \mathbf{l} imes \mathbf{F}_{g},$

where I is the length vector of the pendulum and  $\mathbf{F}_{\mathrm{g}}$  is the force due to gravity.

For now just consider the magnitude of the torque on the pendulum.

 $|\boldsymbol{\tau}| = -mg\ell\sin heta,$ 

where *m* is the mass of the pendulum, *g* is the acceleration due to gravity, *l* is the length of the pendulum, and  $\theta$  is the angle between the length vector and the force due to gravity.

Next rewrite the angular momentum.

$$\mathbf{L} = \mathbf{r} imes \mathbf{p} = m\mathbf{r} imes (oldsymbol{\omega} imes \mathbf{r}).$$

Again just consider the magnitude of the angular momentum.

$$|\mathbf{L}|=mr^{2}\omega=m\ell^{2}rac{d heta}{dt}.$$

and its time derivative

$$rac{d}{dt} |{f L}| = m \ell^2 rac{d^2 heta}{dt^2},$$

According to  $\tau = \frac{d\mathbf{L}}{dt}$ , we can get by comparing the magnitudes

$$-mg\ell\sin heta=m\ell^2rac{d^2 heta}{dt^2},$$

thus:

$$rac{d^2 heta}{dt^2}+rac{g}{\ell}\sin heta=0,$$

which is the same result as obtained through force analysis.





**Figure 2.** Trigonometry of a simple gravity pendulum.

It can also be obtained via the conservation of mechanical energy principle: any object falling a vertical distance h would acquire kinetic energy equal to that which it lost to the fall. In other words, gravitational potential energy is converted into kinetic energy. Change in potential energy is given by

$$\Delta U = mgh.$$

The change in kinetic energy (body started from rest) is given by

$$\Delta K = rac{1}{2}mv^2.$$

Since no energy is lost, the gain in one must be equal to the loss in the other

$$\frac{1}{2}mv^2 = mgh.$$

The change in velocity for a given change in height can be expressed as

$$v=\sqrt{2gh}.$$

Using the arc length formula above, this equation can be rewritten in terms of  $\frac{d\theta}{dt}$ :

$$egin{aligned} v &= \ell rac{d heta}{dt} = \sqrt{2gh}, & ext{so}\ rac{d heta}{dt} &= rac{\sqrt{2gh}}{l}, \end{aligned}$$

where *h* is the vertical distance the pendulum fell. Look at Figure 2, which presents the trigonometry of a simple pendulum. If the pendulum starts its swing from some initial angle  $\theta_0$ , then  $y_0$ , the vertical distance from the screw, is given by

$$y_0 = \ell \cos heta_0.$$

Similarly, for  $y_1$ , we have

$$y_1 = \ell \cos \theta.$$

Then h is the difference of the two

$$h = \ell \left( \cos \theta - \cos \theta_0 \right).$$

In terms of  $\frac{d\theta}{dt}$  gives

(Eq. 2)

This equation is known as the *first integral of motion*, it gives the velocity in terms of the location and includes an integration constant related to the initial displacement

 $(\theta_0)$ . We can differentiate, by applying the chain rule, with respect to time to get the acceleration

$$\begin{split} \frac{d}{dt}\frac{d\theta}{dt} &= \frac{d}{dt}\sqrt{\frac{2g}{\ell}\left(\cos\theta - \cos\theta_{0}\right)},\\ \frac{d^{2}\theta}{dt^{2}} &= \frac{1}{2}\frac{-\frac{2g}{\ell}\sin\theta}{\sqrt{\frac{2g}{\ell}\left(\cos\theta - \cos\theta_{0}\right)}}\frac{d\theta}{dt}\\ &= \frac{1}{2}\frac{-\frac{2g}{\ell}\sin\theta}{\sqrt{\frac{2g}{\ell}\left(\cos\theta - \cos\theta_{0}\right)}}\sqrt{\frac{2g}{\ell}\left(\cos\theta - \cos\theta_{0}\right)} = -\frac{g}{\ell}\sin\theta,\\ \frac{d^{2}\theta}{dt^{2}} &+ \frac{g}{\ell}\sin\theta = 0, \end{split}$$

which is the same result as obtained through force analysis.

## Small-angle approximation



Small-angle approximation for the sine function: For  $\theta \approx 0$  we find  $\sin \theta \approx \theta$ .

The differential equation given above is not easily solved, and there is no solution that can be written in terms of elementary functions. However, adding a restriction to the size of the oscillation's amplitude gives a form whose solution can be easily obtained. If it is assumed that the angle is much less than 1 radian (often cited as less than 0.1 radians, about 6°), or

### $heta\ll 1,$

then substituting for  $\sin \theta$  into **Eq. 1** using the small-angle approximation,

$$\sin\theta \approx heta,$$

yields the equation for a harmonic oscillator,

$$rac{d^2 heta}{dt^2}+rac{g}{\ell} heta=0$$

The error due to the approximation is of order  $\theta^3$  (from the Taylor expansion for sin  $\theta$ ).

Let the starting angle be  $\theta_{0}$ . If it is assumed that the pendulum is released with zero angular velocity, the solution becomes

$$heta(t) = heta_0 \cos\!\left(\sqrt{rac{g}{\ell}}\,t
ight) \qquad heta_0 \ll 1.$$

The motion is simple harmonic motion where  $\theta_0$  is the amplitude of the oscillation (that is, the maximum angle between the rod of the pendulum and the vertical). The corresponding approximate period of the motion is then

$$T_0=2\pi\sqrt{rac{\ell}{g}} \qquad \qquad heta_0\ll 1$$

which is known as Christiaan Huygens's law for the period. Note that under the small-angle approximation, the period is independent of the amplitude  $\theta_0$ ; this is the property of isochronism that Galileo discovered.

### Rule of thumb for pendulum length

$$T_0=2\pi\sqrt{rac{\ell}{g}}$$

gives

$$\ell=rac{g}{\pi^2}rac{T_0^2}{4}.$$

If SI units are used (i.e. measure in metres and seconds), and assuming the measurement is taking place on the Earth's surface, then  $g \approx 9.81 \text{ m/s}^2$ , and  $\frac{g}{\pi^2} \approx 1 \text{ m/s}^2$  (0.994 is the approximation to 3 decimal places).

Therefore, relatively reasonable approximations for the length and period are:



where  $T_0$  is the number of seconds between *two* beats (one beat for each side of the swing), and *l* is measured in metres.

## Arbitrary-amplitude period



**Figure 3.** Deviation of the "true" period of a pendulum from the small-angle approximation of the period. "True" value was obtained numerically evaluating the elliptic integral.



**Figure 4.** Relative errors using the power series for the period.



**Figure 5.** Potential energy and phase portrait of a simple pendulum. Note that the *x*-axis, being angle, wraps onto itself after every  $2\pi$  radians.

For amplitudes beyond the small angle approximation, one can compute the exact period by first inverting the equation for the angular velocity obtained from the energy method (**Eq. 2**),

$$rac{dt}{d heta} = \sqrt{rac{\ell}{2g}} rac{1}{\sqrt{\cos heta - \cos heta_0}}$$

and then integrating over one complete cycle,

$$T=t( heta_0 o 0 o - heta_0 o 0 o heta_0),$$

or twice the half-cycle

$$T=2t( heta_0
ightarrow 0
ightarrow - heta_0),$$

or four times the quarter-cycle

$$T=4t( heta_0
ightarrow 0),$$

which leads to

$$T=4\sqrt{rac{\ell}{2g}}\int_{0}^{ heta_{0}}rac{d heta}{\sqrt{\cos heta-\cos heta_{0}}}$$

Note that this integral diverges as  $\theta_0$  approaches the vertical

$$\lim_{ heta_0
ightarrow\pi}T=\infty,$$

so that a pendulum with just the right energy to go vertical will never actually get there. (Conversely, a pendulum close to its maximum can take an arbitrarily long time to fall down.)

This integral can be rewritten in terms of elliptic integrals as

$$T=4\sqrt{rac{\ell}{g}}F\left(rac{\pi}{2},\sinrac{ heta_{0}}{2}
ight)$$

where F is the incomplete elliptic integral of the first kind defined by

$$F(arphi,k) = \int_0^arphi rac{du}{\sqrt{1-k^2\sin^2 u}} \, .$$

Or more concisely by the substitution

$$\sin u = rac{\sin rac{ heta}{2}}{\sin rac{ heta_0}{2}}$$

expressing  $\theta$  in terms of u,

$$T=rac{2T_0}{\pi}K(k),$$
 where  $k=\sinrac{ heta_0}{2}.$  Eq. 3

Here *K* is the complete elliptic integral of the first kind defined by

$$K(k)=F\left(rac{\pi}{2},k
ight)=\int_{0}^{rac{\pi}{2}}rac{du}{\sqrt{1-k^{2}\sin^{2}u}}\,.$$

For comparison of the approximation to the full solution, consider the period of a pendulum of length 1 m on Earth ( $g = 9.806\ 65\ m/s^2$ ) at initial angle 10 degrees is

$$4\sqrt{rac{1\ \mathrm{m}}{g}}\ K\left(\sinrac{10^\circ}{2}
ight)pprox 2.0102\ \mathrm{s}.$$

The linear approximation gives

$$2\pi\sqrt{rac{1~ ext{m}}{g}}pprox 2.0064~ ext{s}.$$

The difference between the two values, less than 0.2%, is much less than that caused by the variation of g with geographical location.

From here there are many ways to proceed to calculate the elliptic integral.

### Legendre polynomial solution for the elliptic integral

Given Eq. 3 and the Legendre polynomial solution for the elliptic integral:

$$K(k) = rac{\pi}{2} \sum_{n=0}^{\infty} \left( rac{(2n-1)!!}{(2n)!!} k^n 
ight)^2$$

where *n*!! denotes the double factorial, an exact solution to the period of a simple pendulum is:

$$egin{aligned} T&=2\pi\sqrt{rac{\ell}{g}}\left(1+\left(rac{1}{2}
ight)^2\sin^2rac{ heta_0}{2}+\left(rac{1\cdot3}{2\cdot4}
ight)^2\sin^4rac{ heta_0}{2}+\left(rac{1\cdot3\cdot5}{2\cdot4\cdot6}
ight)^2\sin^6rac{ heta_0}{2}+\cdots
ight)\ &=2\pi\sqrt{rac{\ell}{g}}\cdot\sum_{n=0}^\infty\left(\left(rac{(2n)!}{(2^n\cdot n!)^2}
ight)^2\cdot\sin^{2n}rac{ heta_0}{2}
ight). \end{aligned}$$

Figure 4 shows the relative errors using the power series.  $T_0$  is the linear approximation, and  $T_2$  to  $T_{10}$  include respectively the terms up to the 2nd to the 10th powers.

### Power series solution for the elliptic integral

Another formulation of the above solution can be found if the following Maclaurin series:

$$\sinrac{ heta_0}{2}=rac{1}{2} heta_0-rac{1}{48} heta_0^3+rac{1}{3\,840} heta_0^5-rac{1}{645\,120} heta_0^7+\cdots.$$

is used in the Legendre polynomial solution above. The resulting power series is:<sup>[5]</sup>

$$T = 2\pi \sqrt{\frac{\ell}{g}} \left( 1 + \frac{1}{16} \theta_0^2 + \frac{11}{3072} \theta_0^4 + \frac{173}{737280} \theta_0^6 + \frac{22\,931}{1\,321\,205\,760} \theta_0^8 + \frac{1\,319\,183}{951\,268\,147\,200} \theta_0^{10} + \frac{233\,526\,463}{2\,009\,078\,326\,886\,400} \theta_0^{12} + \cdots \right),$$

more fractions available in the On-Line Encyclopedia of Integer Sequences with OEIS: A223067 having the numerators and OEIS: A223068 having the denominators.

### Arithmetic-geometric mean solution for elliptic integral

Given Eq. 3 and the arithmetic–geometric mean solution of the elliptic integral:

$$K(k)=rac{\pi}{2M(1-k,1+k)},$$

where M(x,y) is the arithmetic-geometric mean of x and y.

This yields an alternative and faster-converging formula for the period: [6][7][8]

$$T=rac{2\pi}{M\left(1,\cosrac{ heta_0}{2}
ight)}\sqrt{rac{\ell}{g}}.$$

The first iteration of this algorithm gives

$$T_1=rac{2T_0}{1+\cosrac{ heta_0}{2}}.$$

This approximation has the relative error of less than 1% for angles up to 96.11 degrees.<sup>[6]</sup> Since  $\frac{1}{2}\left(1+\cos\left(\frac{\theta_0}{2}\right)\right)=\cos^2\frac{\theta_0}{4}$ , the expression can be written more concisely as

$$T_1=T_0\sec^2rac{ heta_0}{4}.$$

The second order expansion of  $\sec^2( heta_0/4)$  reduces to  $T pprox T_0\left(1+rac{ heta_0^2}{16}
ight).$ 

A second iteration of this algorithm gives

$$T_2 = rac{4T_0}{1+\cosrac{ heta_0}{2}+2\sqrt{\cosrac{ heta_0}{2}}} = rac{4T_0}{\left(1+\sqrt{\cosrac{ heta_0}{2}}
ight)^2}$$

This second approximation has a relative error of less than 1% for angles up to 163.10 degrees.<sup>[6]</sup>

## Approximate formulae for the nonlinear pendulum period

Though the exact period T can be determined, for any finite amplitude  $\theta_0 < \pi$  rad, by evaluating the corresponding complete elliptic integral K(k), where  $k \equiv \sin(\theta_0/2)$ , this is often avoided in applications because it is not possible to express this integral in a closed form in terms of elementary functions. This has made way for research on simple approximate formulae for the increase of the pendulum period with amplitude (useful in introductory physics labs, classical mechanics, electromagnetism, acoustics, electronics, superconductivity, etc.<sup>[9]</sup> The approximate formulae formula

- 'Not so large-angle' formulae, i.e. those yielding good estimates for amplitudes below  $\pi/2$  rad (a natural limit for a bob on the end of a flexible string), though the deviation with respect to the exact period increases monotonically with amplitude, being unsuitable for amplitudes near to  $\pi$  rad. One of the simplest formulae found in literature is the following one by Lima (2006):  $T \approx -T_0 \frac{\ln a}{1-a}$ , where  $a \equiv \cos(\theta_0/2)$ .<sup>[10]</sup>
- 'Very large-angle' formulae, i.e. those which approximate the exact period asymptotically for amplitudes near to  $\pi$ rad, with an error that increases monotonically for smaller amplitudes (i.e., unsuitable for small amplitudes). One of the better such formulae is that by Cromer, namely:<sup>[11]</sup>  $T \approx \frac{2}{\pi} T_0 \ln (4/a)$ .

Of course, the increase of T with amplitude is more apparent when  $\pi/2 < \theta_0 < \pi$ , as has been observed in many experiments using either a rigid rod or a disc.<sup>[12]</sup> As accurate timers and sensors are currently available even in introductory physics labs, the experimental errors found in 'very large-angle' experiments are already small enough for a comparison with the exact period and a very good agreement between theory and experiments in which friction is negligible has been found. Since this activity has been encouraged by many instructors, a simple approximate formula for the pendulum period valid for all possible amplitudes, to which experimental data could be compared, was sought. In 2008, Lima derived a weighted-average formula with this characteristic:<sup>[9]</sup>

$$Tpprox rac{r\,a^2\,T_{
m Lima}+k^2\,T_{
m Cromer}}{r\,a^2+k^2},$$

where r=7.17, which presents a maximum error of only 0.6% (at  $heta_0=95^\circ$ ).

## Arbitrary-amplitude angular displacement Fourier series

The Fourier series expansion of  $\theta(t)$  is given by

$$heta(t) = 8\sum_{n\geq 1 ext{ odd}} rac{(-1)^{\lfloor n/2 
floor}}{n} rac{q^{n/2}}{1+q^n} \cos(n\omega t)$$

where q is the elliptic nome,  $q=\exp(-\pi K'/K)$ , and  $\omega=2\pi/T$  the angular frequency. If one defines

$$\epsilon = rac{1-\sqrt{\cosrac{ heta_0}{2}}}{2+2\sqrt{\cosrac{ heta_0}{2}}}$$

q can be approximated using the expansion

 $q = \varepsilon + 2\varepsilon^5 + 15\varepsilon^9 + 150\varepsilon^{13} + 1707\varepsilon^{17} + 20910\varepsilon^{21} + \cdots$ (see OEIS: A002103). Note that for  $\theta_0 < \pi$  we have  $\varepsilon < \frac{1}{2}$ , thus the approximation is applicable even for large amplitudes.

## Examples

The animations below depict the motion of a simple (frictionless) pendulum with increasing amounts of initial displacement of the bob, or equivalently increasing initial velocity. The small graph above each pendulum is the corresponding phase plane diagram; the horizontal axis is displacement and the vertical axis is velocity. With a large enough initial velocity the pendulum does not oscillate back and forth but rotates completely around the pivot.



### Initial angle of 0°, a stable equilibrium



### Initial angle of 45°

•×

Initial angle of 90°



Initial angle of 135°



Initial angle of 170°



Initial angle of 180°, unstable equilibrium



Pendulum with just barely enough energy for a full swing



Pendulum with enough energy for a full swing

# Compound pendulum

A **compound pendulum** (or **physical pendulum**) is one where the rod is not massless, and may have extended size; that is, an arbitrarily shaped rigid body swinging by a pivot. In this case the pendulum's period depends on its moment of inertia *I* around the pivot point.

The equation of torque gives:

$$au = I lpha$$

where:

- $\alpha$  is the angular acceleration.
- $\tau$  is the torque

The torque is generated by gravity so:

$$au = -mgL\sin heta$$

where:

- *m* is the mass of the body
- L is the distance from the pivot to the object's center of mass
- $\theta$  is the angle from the vertical

Hence, under the small-angle approximation  $\sin \theta \approx \theta$ ,

$$lpha = \ddot{ heta} pprox - rac{mgL heta}{I}$$

where *I* is the moment of inertia of the body about the pivot point.

The expression for  $\alpha$  is of the same form as the conventional simple pendulum and gives a period of<sup>[2]</sup>

$$T=2\pi\sqrt{rac{I}{mgL}}$$

And a frequency of

$$f=rac{1}{T}=rac{1}{2\pi}\sqrt{rac{mgL}{I}}$$

If the initial angle is taken into consideration (for large amplitudes), then the expression for lpha becomes:

$$lpha = \ddot{ heta} = -rac{mgL\sin heta}{I}$$

and gives a period of:

$$T = 4\,{
m K}igg({
m sin}^2\,{ heta_0\over 2}igg)\sqrt{{I\over mgL}}$$

where  $\theta_0$  is the maximum angle of oscillation (with respect to the vertical) and K(k) is the complete elliptic integral of the first kind.

## Physical interpretation of the imaginary period

The Jacobian elliptic function that expresses the position of a pendulum as a function of time is a doubly periodic function with a real period and an imaginary period. The real period is, of course, the time it takes the pendulum to go through one full cycle. Paul Appell pointed out a physical interpretation of the imaginary period:<sup>[13]</sup> if  $\theta_0$  is the maximum angle of one pendulum and  $180^\circ - \theta_0$  is the maximum angle of another, then the real period of each is the magnitude of the imaginary period of the other.

## Coupled pendula



Two identical simple pendulums coupled via a spring connecting the bobs.

**Coupled pendulums** can affect each other's motion, either through a direction connection (such as a spring connecting the bobs) or through motions in a supporting structure (such as a tabletop). The equations of motion for two identical simple pendulums coupled by a spring connecting the bobs can be obtained using Lagrangian Mechanics.

The kinetic energy of the system is:

$$E_{\mathrm{K}}=rac{1}{2}mL^{2}\left(\dot{ heta}_{1}^{2}+\dot{ heta}_{2}^{2}
ight)$$

where m is the mass of the bobs, L is the length of the strings, and  $\theta_1$ ,  $\theta_2$  are the angular displacements of the two bobs from equilibrium.

The potential energy of the system is:

$$E_{\mathrm{p}}=mgL(2-\cos heta_{1}-\cos heta_{2})+rac{1}{2}kL^{2}( heta_{2}- heta_{1})^{2}$$

where g is the gravitational acceleration, and k is the spring constant. The displacement  $L(\theta_2 - \theta_1)$  of the spring from its equilibrium position assumes the small angle approximation.

The Lagrangian is then

$$\mathcal{L} = rac{1}{2}mL^2\left({\dot heta}_1^2 + {\dot heta}_2^2
ight) - mgL(2 - \cos heta_1 - \cos heta_2) - rac{1}{2}kL^2( heta_2 - heta_1)^2$$

which leads to the following set of coupled differential equations:

$$egin{aligned} \ddot{ heta}_1 + rac{g}{L}\sin heta_1 + rac{\kappa}{m}( heta_1 - heta_2) &= 0 \ \ddot{ heta}_2 + rac{g}{L}\sin heta_2 - rac{k}{m}( heta_1 - heta_2) &= 0 \end{aligned}$$

Adding and subtracting these two equations in turn, and applying the small angle approximation, gives two harmonic oscillator equations in the variables  $\theta_1 + \theta_2$  and  $\theta_1 - \theta_2$ :

$$\ddot{ heta}_1+\ddot{ heta}_2+rac{g}{L}( heta_1+ heta_2)=0 \ \ddot{ heta}_1-\ddot{ heta}_2+\left(rac{g}{L}+2rac{k}{m}
ight)( heta_1- heta_2)=0$$

with the corresponding solutions

$$egin{aligned} heta_1+ heta_2&=A\cos(\omega_1t+lpha)\ heta_1- heta_2&=B\cos(\omega_2t+eta) \end{aligned}$$

where

$$\omega_1 = \sqrt{rac{g}{L}} \ \omega_2 = \sqrt{rac{g}{L}+2rac{k}{m}}$$

and A, B,  $\alpha$ ,  $\beta$  are constants of integration.

Expressing the solutions in terms of  $\theta_1$  and  $\theta_2$  alone:

$$egin{aligned} heta_1&=rac{1}{2}A\cos(\omega_1t+lpha)+rac{1}{2}B\cos(\omega_2t+eta)\ heta_2&=rac{1}{2}A\cos(\omega_1t+lpha)-rac{1}{2}B\cos(\omega_2t+eta) \end{aligned}$$

If the bobs are not given an initial push, then the condition  $\dot{\theta}_1(0) = \dot{\theta}_2(0) = 0$  requires  $\alpha = \beta = 0$ , which gives (after some rearranging):

$$egin{aligned} A&= heta_1(0)+ heta_2(0)\ B&= heta_1(0)- heta_2(0) \end{aligned}$$

### See also

- Blackburn pendulum
- Conical pendulum
- Cycloidal pendulum
- Double pendulum
- Inverted pendulum
- Kapitza's pendulum
- Rayleigh–Lorentz pendulum
- Spring pendulum
- Mathieu function
- Pendulum equations (software)

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### External links

• Mathworld article on Mathieu Function (http://mathworld.wolfram.com/MathieuFunction.html)